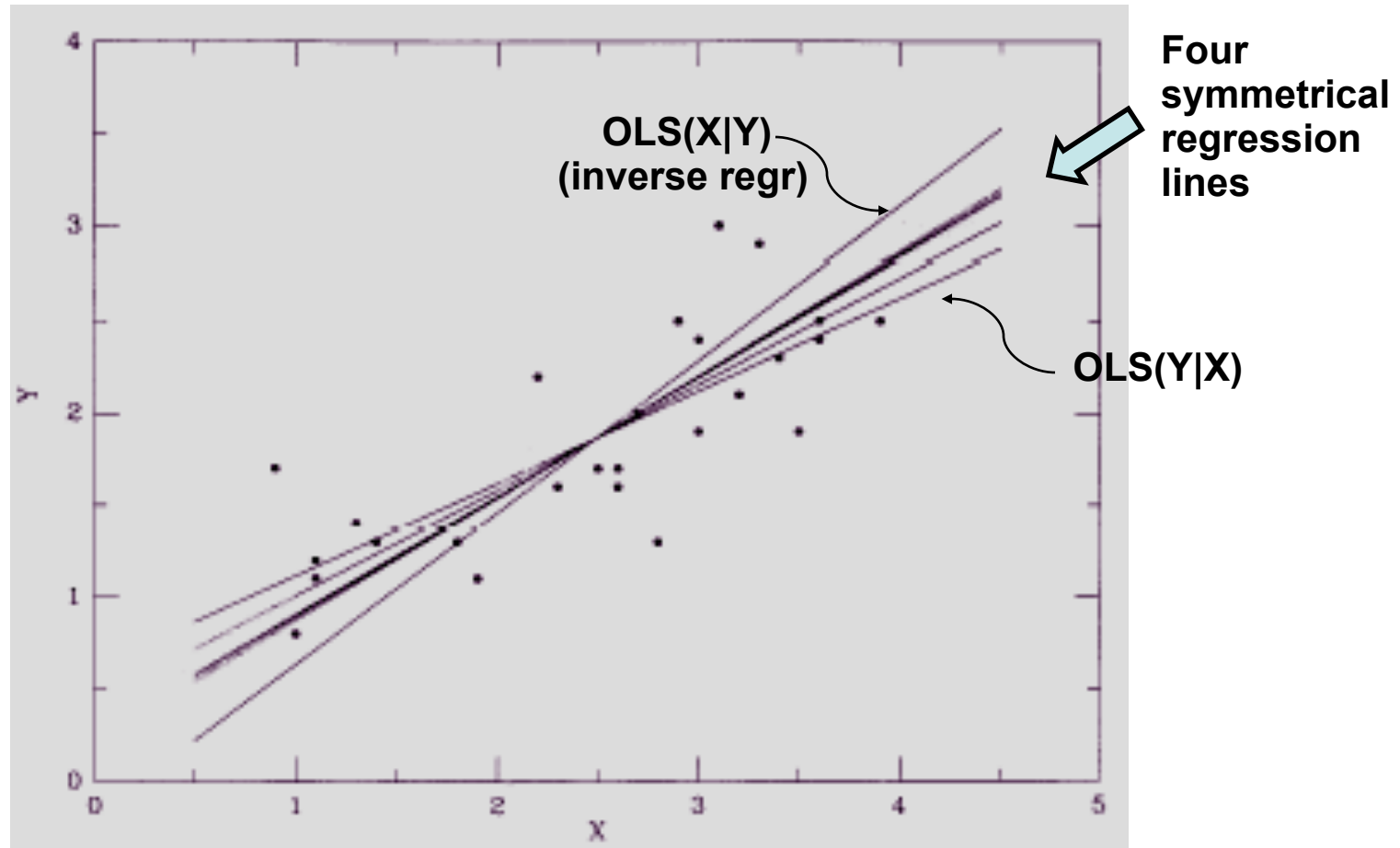


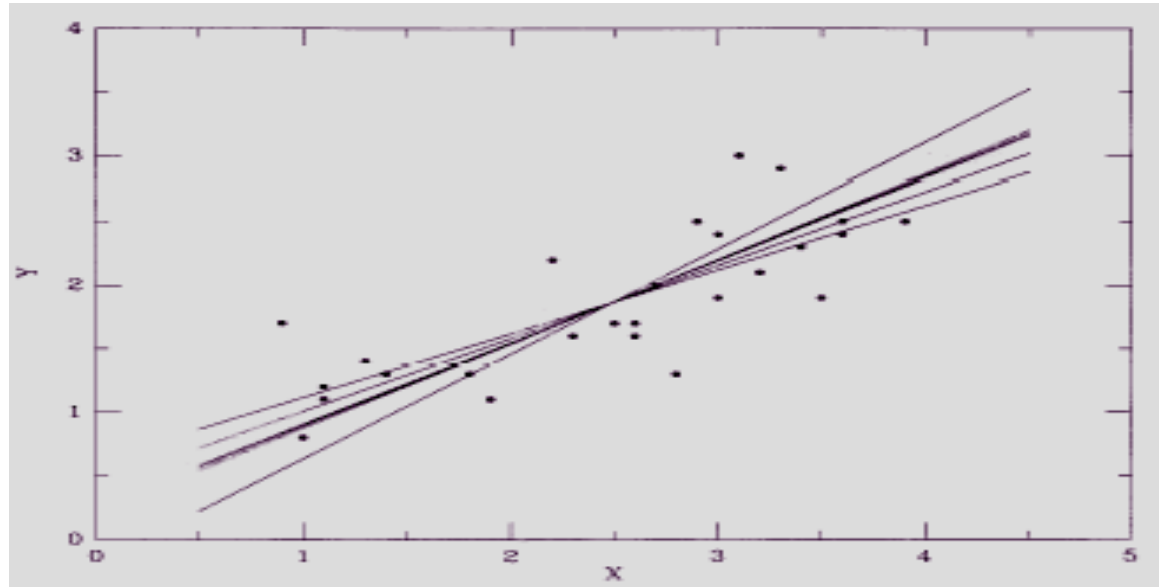
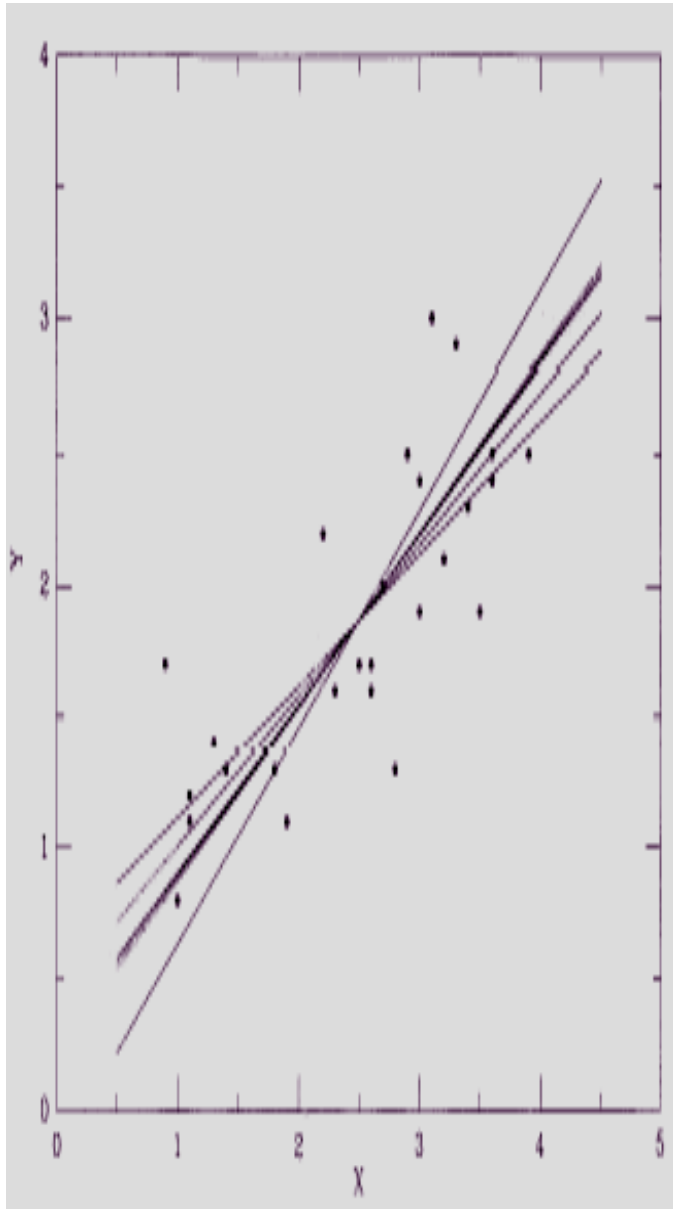
# Linear regression issues in astronomy

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# Structural regression

Seeking the intrinsic relationship between two properties without specifying 'dependent' and 'independent' variables





**Shrinking and Stretching**  
Slope depends on the units

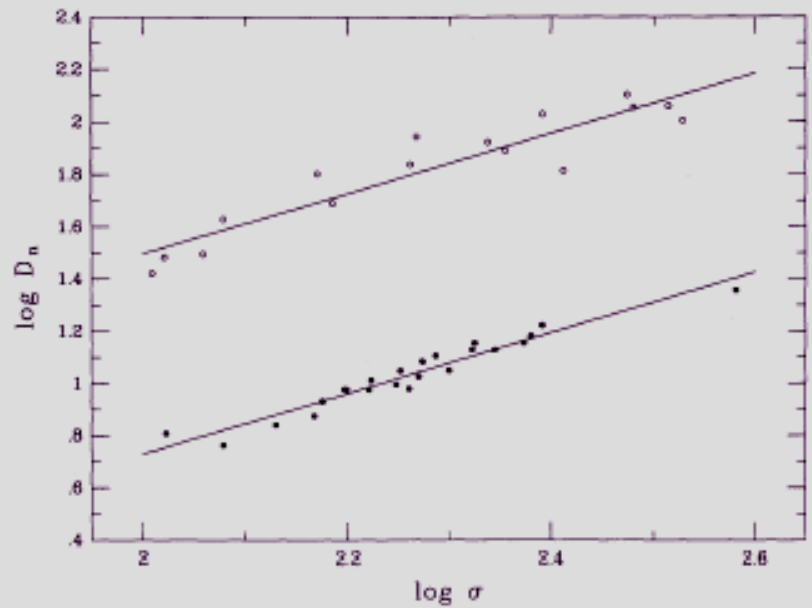
# Analytical formulae for slopes of the 6 OLS lines

TABLE 1  
LINEAR REGRESSION FORMULAE FOR SLOPES

Method	Expression for Slope	Estimate of the Variance of the Slope $\widehat{\text{Var}}(\beta_i)$
OLS(X   Y) .....	$\beta_1 = \frac{S_{xy}}{S_{xx}}$	$\frac{1}{S_{xx}^2} \left[ \sum_{i=1}^n (x_i - \bar{x})^2 (y_i - \beta_1 x_i - \bar{y} + \beta_1 \bar{x})^2 \right]$
OLS(Y   X) .....	$\beta_2 = \frac{S_{yy}}{S_{xy}}$	$\frac{1}{S_{xy}^2} \left[ \sum_{i=1}^n (y_i - \bar{y})^2 (y_i - \beta_2 x_i - \bar{y} + \beta_2 \bar{x})^2 \right]$
OLS bisector .....	$\beta_3 = (\beta_1 + \beta_2)^{-1} [\beta_1 \beta_2 - 1 + \sqrt{(1 + \beta_1^2)(1 + \beta_2^2)}]$	$\frac{\beta_3^2}{(\beta_1 + \beta_2)^2 (1 + \beta_1^2)(1 + \beta_2^2)} [(1 + \beta_2^2)^2 \widehat{\text{Var}}(\beta_1) + 2(1 + \beta_1^2)(1 + \beta_2^2) \widehat{\text{Cov}}(\beta_1, \beta_2) + (1 + \beta_1^2)^2 \widehat{\text{Var}}(\beta_2)]$
Orthogonal regression .....	$\beta_4 = \frac{1}{2} [(\beta_2 - \beta_1^{-1}) + \text{Sign}(S_{xy}) \sqrt{4 + (\beta_2 - \beta_1^{-1})^2}]$	$\frac{\beta_4^2}{4\beta_1^2 + (\beta_1 \beta_2 - 1)^2} [\beta_1^{-2} \widehat{\text{Var}}(\beta_1) + 2 \widehat{\text{Cov}}(\beta_1, \beta_2) + \beta_1^2 \widehat{\text{Var}}(\beta_2)]$
Reduced major-axis .....	$\beta_5 = \text{Sign}(S_{xy}) (\beta_1 \beta_2)^{1/2}$	$\frac{1}{4} \left[ \frac{\beta_2}{\beta_1} \widehat{\text{Var}}(\beta_1) + 2 \widehat{\text{Cov}}(\beta_1, \beta_2) + \frac{\beta_1}{\beta_2} \widehat{\text{Var}}(\beta_2) \right]$

NOTE.—An estimate of covariance term is given by:

$$\widehat{\text{Cov}}(\beta_1, \beta_2) = (\beta_1 S_{xx}^2)^{-1} \left\{ \sum_{i=1}^n (x_i - \bar{x})(y_i - \bar{y}) [y_i - \bar{y} - \beta_1(x_i - \bar{x})][y_i - \bar{y} - \beta_2(x_i - \bar{x})] \right\}.$$

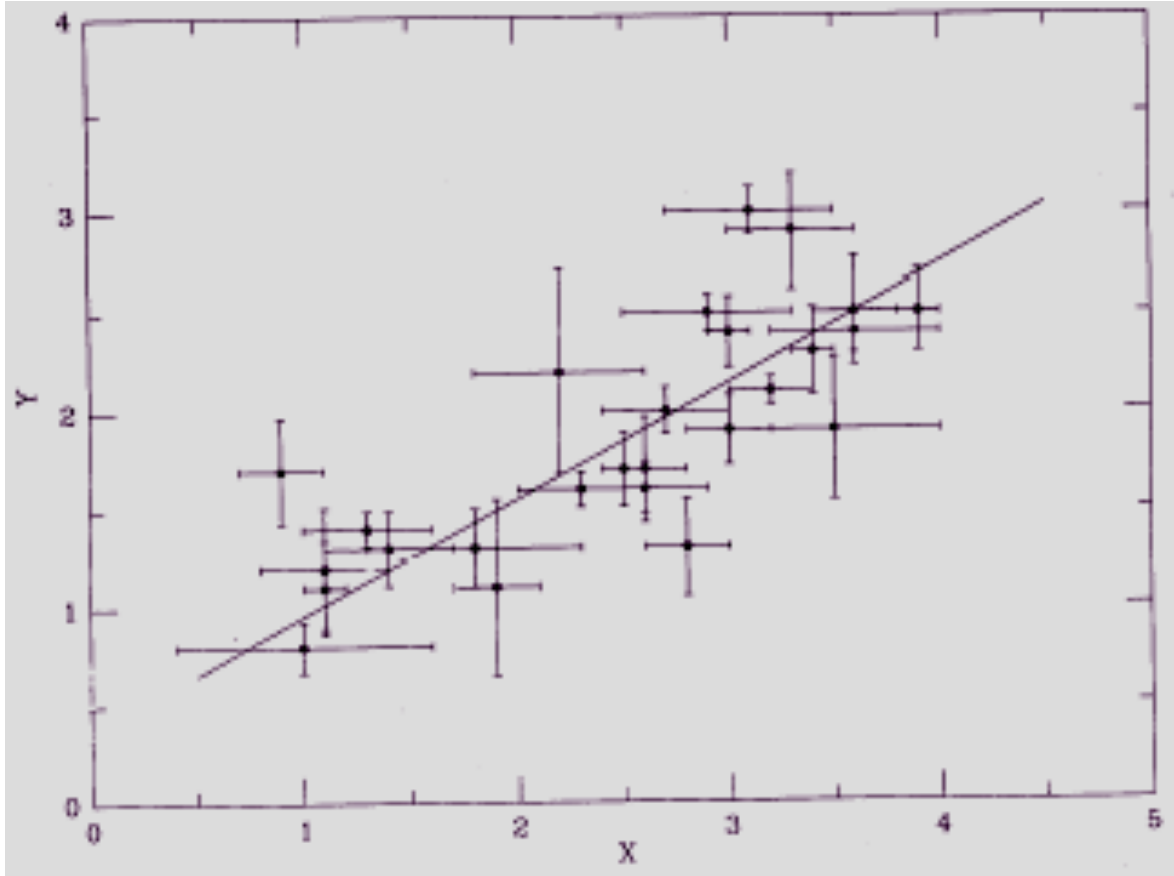


**Example: Faber-Jackson relation between diameter and stellar velocity dispersion of elliptical galaxies**

**TABLE 4**  
REGRESSIONS FOR COMA AND VIRGO  $\log D_n$  VERSUS  $\log \sigma^*$

METHOD (1)	ASYMPTOTIC FORMULAE		BOOTSTRAP SLOPE (4)	JACKKNIFE SLOPE (5)
	Intercept (2)	Slope (3)		
23 Coma Ellipticals				
OLS(Y   X) .....	$-1.595 \pm 0.186$	$1.162 \pm 0.082$	$1.186 \pm 0.094$	$1.164 \pm 0.111$
OLS(X   Y) .....	$-1.765 \pm 0.216$	$1.238 \pm 0.096$	$1.261 \pm 0.104$	$1.239 \pm 0.128$
OLS bisector .....	$-1.678 \pm 0.200$	$1.199 \pm 0.088$	$1.223 \pm 0.099$	$1.201 \pm 0.119$
Orthogonal .....	$-1.694 \pm 0.209$	$1.206 \pm 0.092$	$1.231 \pm 0.102$	$1.208 \pm 0.124$
Reduced major axis .....	$-1.679 \pm 0.200$	$1.199 \pm 0.088$	$1.223 \pm 0.099$	$1.201 \pm 0.119$
OLS mean .....	$-1.680 \pm 0.200$	$1.200 \pm 0.088$	$1.224 \pm 0.099$	$1.201 \pm 0.119$
16 Virgo Ellipticals				
OLS(Y   X) .....	$-0.790 \pm 0.230$	$1.144 \pm 0.101$	$1.143 \pm 0.127$	$1.114 \pm 0.118$
OLS(X   Y) .....	$-1.183 \pm 0.180$	$1.316 \pm 0.082$	$1.322 \pm 0.132$	$1.316 \pm 0.093$
OLS bisector .....	$-0.978 \pm 0.190$	$1.227 \pm 0.085$	$1.227 \pm 0.107$	$1.226 \pm 0.099$
Orthogonal .....	$-1.021 \pm 0.198$	$1.245 \pm 0.089$	$1.246 \pm 0.121$	$1.245 \pm 0.104$
Reduced major axis .....	$-0.979 \pm 0.190$	$1.227 \pm 0.085$	$1.228 \pm 0.108$	$1.227 \pm 0.099$
OLS mean .....	$-0.986 \pm 0.188$	$1.230 \pm 0.084$	$1.233 \pm 0.110$	$1.230 \pm 0.098$

# Heteroscedastic measurement errors in both variables



**Homoscedastic functional**  
Deeming (Vistas Astr 1968)  
Fuller "Measurement Error  
Models" (1987)

**Heteroscedastic functional**  
York (Can J Phys 1966)  
ODRPACK Boggs et al.  
(ACM Trans Math Soft 1990)

**Heteroscedastic structural**  
BCES (Akritas & Bershadsky  
ApJ 1996)

# Regression with measurement errors and intrinsic scatter

**Y = observed data**

**V = measurement errors**

$$(Y_{1i}, Y_{2i}, V_i), \quad i = 1, \dots, n$$

**X = intrinsic variables**

**e = intrinsic scatter**

$$Y_{1i} = X_{1i} + \epsilon_{1i} \quad \text{and} \quad Y_{2i} = X_{2i} + \epsilon_{2i}$$

**Regression model**

$$X_{2i} = \alpha_1 + \beta_1 X_{1i} + \epsilon_i$$

**BCES slope estimator**

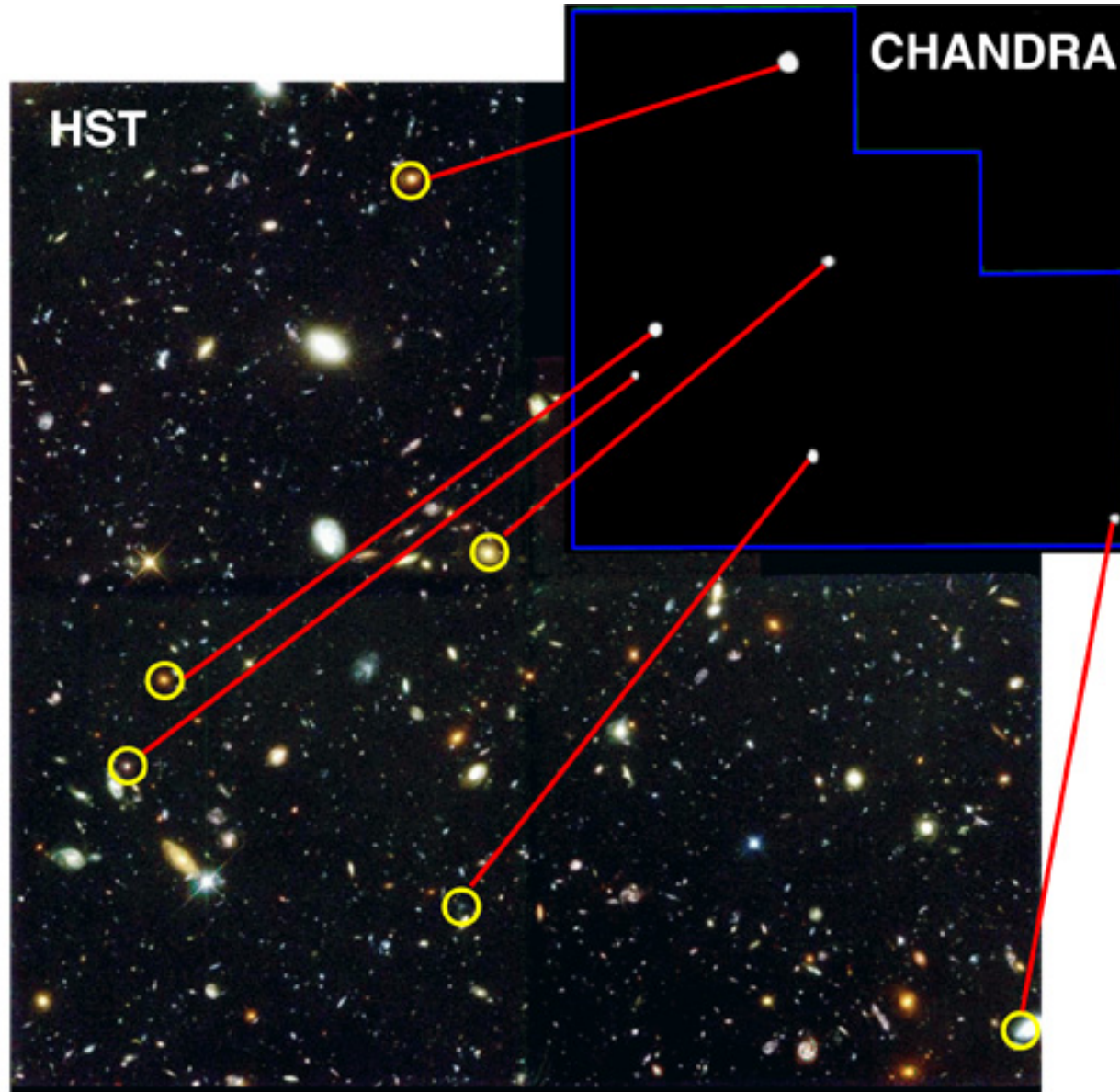
$$\hat{\beta}_1 = \frac{\sum_{i=1}^n (Y_{1i} - \bar{Y}_1)(Y_{2i} - \bar{Y}_2) - \sum_{i=1}^n V_{12,i}}{\sum_{i=1}^n (Y_{1i} - \bar{Y}_1)^2 - \sum_{i=1}^n V_{11,i}}$$

$$\hat{\alpha}_1 = \bar{Y}_2 - \beta_1 \bar{Y}_1.$$

**BCES slope variance**

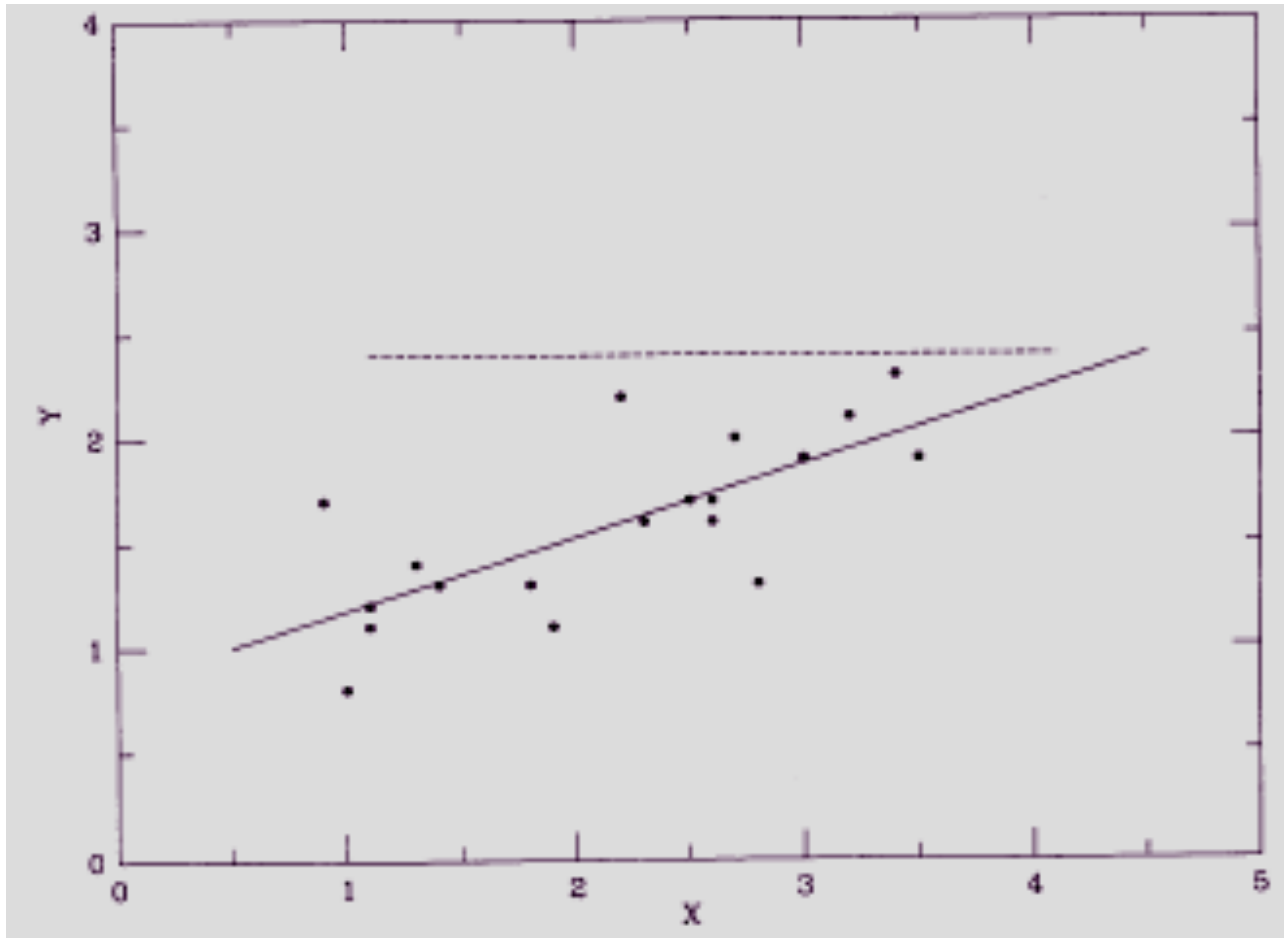
$$\hat{\sigma}_{\beta_1}^2 = n^{-1} \sum_{i=1}^n (\xi_{1i} - \bar{\xi}_1)^2 \quad \xi_{1i} = \frac{[Y_{1i} - E(Y_{1i})](Y_{2i} - \beta_1 Y_{1i} - \alpha_1) + \beta_1 V_{11,i} - V_{12,i}}{V(Y_{1i}) - E(V_{11,i})}$$

# X-ray sources in the HDFN





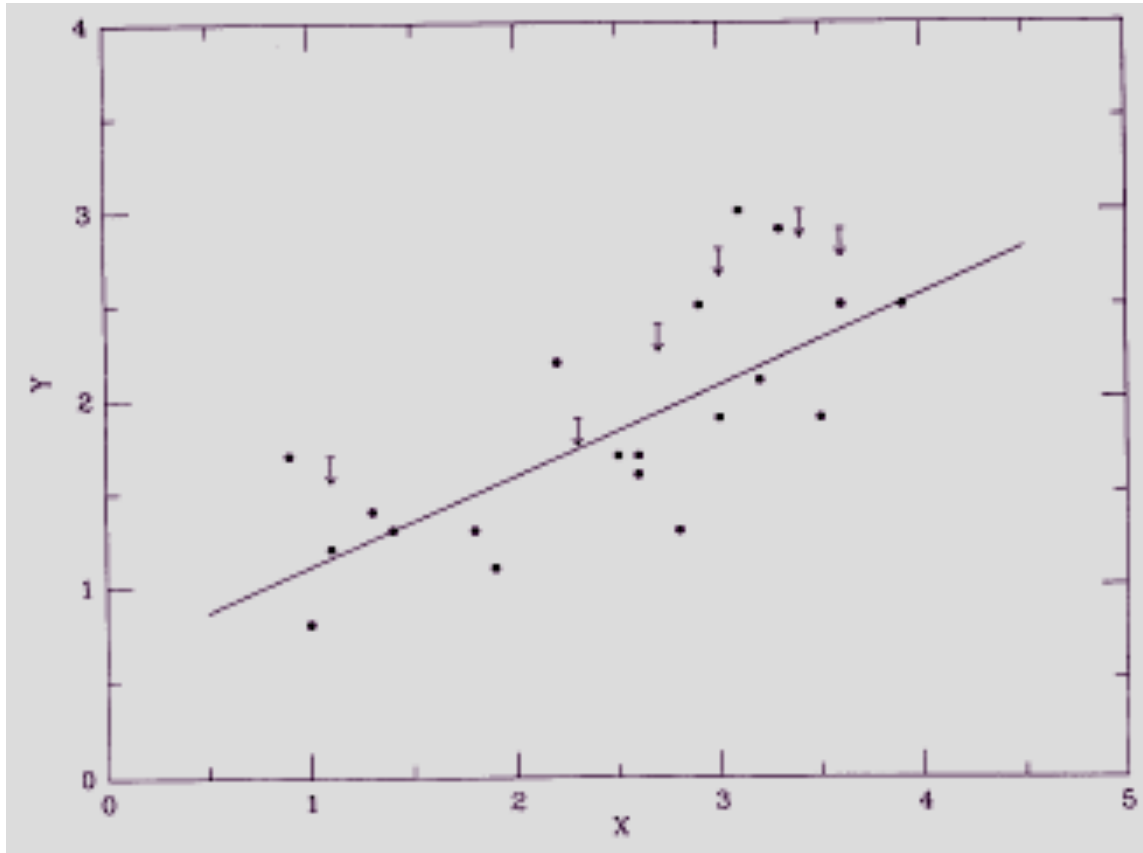
# Truncation due to flux limits



**Econometrics:** Tobit & LIMDEP models (Amemiya, *Advanced econometrics* 1985; Maddala, *Limited-dependent & Quantitative Variables in Econometrics* 1983)

**Astronomy:** Malmquist bias in Hubble diagram (Deeming, *Vistas Astr* 1968, Segal, *PNAS* 1975)

# Censoring due to non-detections



## **Correlation coefficients:**

Generalized Kendall's  $\tau$  (Brown, Hollander & Korwar 1974)

## **Linear regression with normal residuals:**

EM Algorithm (Wolynetz Appl Stat 1979)

## **Linear regression with Kaplan-Meier residuals:**

Buckley & James (Biometrika 1979) Schmitt (ApJ 1985)

**Presented for astronomy by Isobe, Feigelson & Nelson (ApJ 1986)  
Implemented in Astronomy Survival Analysis (ASURV) package**

# Bayesian Treatment of Measurement Errors in Linear Regression

- Errors-in-variables regression model (cf. monograph W. Fuller 1987)

$$\eta_i = \alpha + \beta \xi_i + \varepsilon_i \quad (\text{True relationship})$$

$$x_i = \xi_i + \varepsilon_{x,i} \quad (\text{True variables indirectly observed}$$

$$y_i = \eta_i + \varepsilon_{y,i} \quad (\text{with measurement errors})$$

- $\xi_i$  is modeled as mixtures of Normals

# Mixture of Normals Model

- Model the distribution of  $\xi$  as a mixture of  $K$  Gaussians, assume Gaussian intrinsic scatter and Gaussian measurement errors of known variance
- The model is hierarchically expressed as:

$$\xi_i | \pi, \mu, \tau^2 \sim \sum_{k=1}^K \pi_k N(\mu_k, \tau_k^2)$$

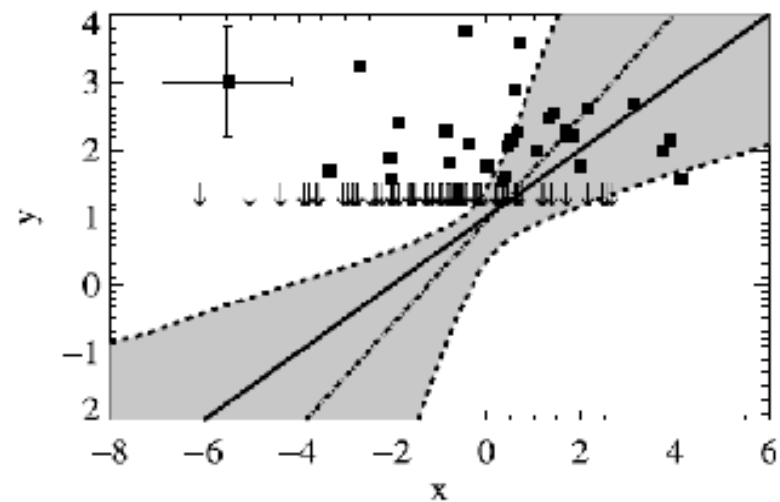
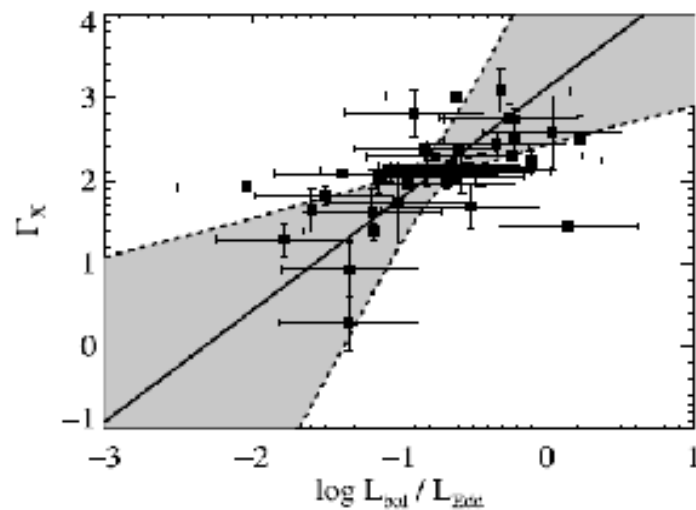
$$\eta_i | \xi_i, \alpha, \beta, \sigma^2 \sim N(\alpha + \beta \xi_i, \sigma^2)$$

$$y_i, x_i | \eta_i, \xi_i \sim N([\eta_i, \xi_i], \Sigma_i)$$

$$\psi = (\pi, \mu, \tau^2), \quad \theta = (\alpha, \beta, \sigma^2), \quad \Sigma_i = \begin{pmatrix} \sigma_{y,i}^2 & \sigma_{xy,i} \\ \sigma_{xy,i} & \sigma_{x,i}^2 \end{pmatrix}$$

Prior distributions are assigned to the parameters  $(\alpha, \beta, \sigma, \tau, \mu)$ ,  
Bayes' Theorem is applied, and posterior distributions are computed  
with Markov chain Monte Carlo techniques.

The method can be applied to censored and truncated regression  
problems, as well as measurement error problems. Performance is  
demonstrably better than earlier de-biased least-squares solutions  
(BCES, FITEXY). IDL code is available.



## Conclusions

**Bivariate linear regression in astronomy can be surprisingly complex. Pay attention to precise question being asked, and details of situation. Several codes available through <http://astrostatistics.psu.edu/statcodes>.**

- **Functional vs. structural regression**
- **Symmetrical vs. dependent regression**
- **Weighting by measurement error**
- **Truncation & censoring due to flux limits**

**Other topics not considered here (some covered later in the Summer School):**

- **Robust & rank regression techniques to treat outliers**
- **Goodness-of-fit, model selection and parsimony**
- **Nonlinear regression**
- **Multivariate regression**

# References

- Isobe, Feigelson, Akritas & Babu, ApJ 364, 105, 1990
- Feigelson and Babu, ApJ 397, 55, 1992
- Brandon Kelly, ApJ 665, 1489, 2007
- W. Fuller 1987